

Spectral Shift Function for the magnetic Schrödinger operators

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Definition of SSF

For two self-adjoint operators H_0 and H on a separable Hilbert space \mathcal{H} , M. G. Kreĭn's **spectral shift function (SSF)** $\xi(\lambda) = \xi(\lambda; H, H_0)$ for the pair (H, H_0) is defined by the trace formula

$$\mathrm{tr} [f(H) - f(H_0)] = \int_{\mathbb{R}} \xi(\lambda) f'(\lambda) d\lambda \left(= - \int_{\mathbb{R}} \xi'(\lambda) f(\lambda) d\lambda \right)$$

for any $f \in C_0^\infty(\mathbb{R})$ (cf. **Kreĭn 1953**). If the spectrum of H and H_0 are included in $(-c, \infty)$ and $(H + cI)^{-m} - (H_0 + cI)^{-m}$ is in the **trace class** for some $c \in \mathbb{R}$, then SSF is defined via the equality

$$\xi(\lambda; H, H_0) = \begin{cases} -\xi((\lambda + c)^{-m}; (H + cI)^{-m}, (H_0 + cI)^{-m}) & (\lambda > -c), \\ 0 & (\lambda \leq -c). \end{cases}$$

SSF for Schrödinger operators

In the case $H_0 = -\Delta$ and $H = -\Delta + V$ on \mathbb{R}^d , a well-known sufficient condition for the existence of SSF is

$$|V(x)| \leq C\langle x \rangle^{-\rho}, \quad \langle x \rangle = (1 + |x|^2)^{1/2} \quad (1)$$

for some $C > 0$ and $\rho > d$. It is also known that **regularized SSF** can be defined under more mild decaying condition depending on the dimension d . At least we always need the short range condition $\rho > 1$.

Nice reviews: **Birman-Yafaev 1992**, **Birman-Pushnitski 1998**, **Yafaev 2007**.

Birman-Kreĭn formula

When $\rho > d$, the scattering matrix $S(\lambda) = S(\lambda; H, H_0)$ exists and $S(\lambda) - I$ is in the trace class. Then the **Birman-Kreĭn formula**

$$\det S(\lambda) = \exp(-2\pi i \xi(\lambda)) \quad (2)$$

holds for almost every $\lambda > 0$ (**Birman-Kreĭn 1962**). If we write the eigenvalues of $S(\lambda)$ as $e^{2i\delta_{\lambda,n}}$ ($n = 1, 2, \dots$), we have

$$\xi(\lambda) = -\frac{1}{\pi} \sum_{n=1}^{\infty} \delta_{\lambda,n}. \quad (3)$$

The RHS of (3) is absolutely summable when $\rho > d$. The number $\delta_{\lambda,n}$ is called the **phase shift** when V is radial, since $\delta_{\lambda,n}$ is just the asymptotic phase shift of some generalized eigenfunction for H .

Magnetic Schrödinger operator

Next we consider the magnetic Schrödinger operator on \mathbb{R}^2

$$H = \left(\frac{1}{i} \nabla - A \right)^2, \quad A = (A_1, A_2).$$

The corresponding magnetic field and the total magnetic flux are

$$B = \text{curl } A = \partial_1 A_2 - \partial_2 A_1, \quad \alpha = \frac{1}{2\pi} \int_{\mathbb{R}^2} B(x) dx.$$

If the vector potential A satisfies

$$|A(x)| + |\text{div } A(x)| \leq C \langle x \rangle^{-\rho}, \quad \rho > 2, \quad (4)$$

then we can also define SSF $\xi(\lambda; H, H_0)$ ($H_0 = -\Delta$) in a similar way. However, (4) never holds when $\alpha \neq 0$, and we cannot define SSF in the ordinary manner.

Main result

Nevertheless, we can define similar quantity **even if $\alpha \neq 0$** , in the following sense.

Theorem 1

Assume the magnetic field B is a real-valued C^1 function on \mathbb{R}^2 such that

$$|B(x)| \leq C \langle x \rangle^{-\rho}, \quad \rho > 3.$$

Let $\alpha = \int_{\mathbb{R}^2} B(x) dx / (2\pi)$ be the total magnetic flux, and H_α be the Schrödinger operator for the Aharonov-Bohm magnetic field

$$H_\alpha = \left(\frac{1}{i} \nabla - A_\alpha \right)^2, \quad A_\alpha = \alpha \left(-\frac{x_2}{|x|^2}, \frac{x_1}{|x|^2} \right)$$

with the regular boundary condition at $x = 0$.

Theorem 1 (continued)

Then, there exists a vector potential A with $\text{curl } A = B$, such that

$$\begin{aligned} & \lim_{R \rightarrow \infty} \text{tr} [\chi_R (f(H) - f(H_0)) \chi_R] \\ &= -\frac{1}{2} \{\alpha\} (1 - \{\alpha\}) f(0) + \int_{\mathbb{R}} \xi(\lambda; H, H_\alpha) f'(\lambda) d\lambda \end{aligned}$$

for every $f \in C_0^\infty(\mathbb{R})$. Here χ_R is the characteristic function of the disc $\{|x| \leq R\}$, $\{\alpha\} = \alpha - [\alpha]$ is the fractional part of α , and $\xi(\lambda; H, H_\alpha)$ is the ordinary SSF for the pair (H, H_α) .

Similar results:

- **Borg 2006 (Ph. D. thesis)** $f = e^{-t\lambda}$, $H = H_\alpha$, with Dirichlet b.c.
- **Tamura 2008** $f' = 0$ near the origin, χ_R is replaced by the smooth cut-off function.

SSF for Aharonov-Bohm magnetic field

The above result is formally interpreted as

$$\begin{aligned}\xi(\lambda; H, H_0) &= \xi(\lambda; H, H_\alpha) + \xi(\lambda; H_\alpha, H_0), \\ \xi(\lambda; H_\alpha, H_0) &= \begin{cases} \frac{1}{2}\{\alpha\}(1 - \{\alpha\}) & (\lambda > 0), \\ 0 & (\lambda \leq 0). \end{cases}\end{aligned}$$

The eigenvalues of the scattering matrix $S(\lambda) = S(\lambda; H_\alpha, H_0)$ are $e^{i\alpha\pi}$ and $e^{-i\alpha\pi}$ (∞ -deg.) (Ruijsenaars 1983, Adami-Teta 1998, Roux-Yafaev 2002). Then Birman-Kreĭn formula becomes

$$\begin{aligned}\xi(\lambda; H_\alpha, H_0) &= \left(-\frac{\{\alpha\}}{2} - \frac{\{\alpha\}}{2} - \frac{\{\alpha\}}{2} - \dots \right) \\ &\quad + \left(\frac{\{\alpha\}}{2} + \frac{\{\alpha\}}{2} + \frac{\{\alpha\}}{2} + \frac{\{\alpha\}}{2} + \dots \right).\end{aligned}$$

This equality does not make sense at all, but it also suggests us there is some cancellation mechanism.

Outline of Proof

The key tool for the proof of Theorem 1 is **Pochhammer's generalized hypergeometric function**

$${}_pF_q(\alpha_1, \dots, \alpha_p; \beta_1, \dots, \beta_q; z) = \sum_{n=0}^{\infty} \frac{\Gamma(\alpha_1 + n) \cdots \Gamma(\alpha_p + n)}{\Gamma(\beta_1 + n) \cdots \Gamma(\beta_q + n)} \frac{z^n}{n!}.$$

Here we obey E. M. Wright's notation. The asymptotic formula for ${}_pF_q$ has been studied from the beginning of 20th century (cf. **Barnes 1907, Wright 1935, 1940, Braaksma 1962, Luke 1969, 1975, ...**). The asymptotic formula consists of **algebraic series** and **exponential series**, whose coefficients can be explicitly calculated (at least by Mathematica).

Outline of Proof

Proposition 2

Let $0 < \alpha < 1$ and $f \in C_0^\infty(\mathbb{R})$. Then, we have

$$\mathrm{tr} [\chi_R (f(H_\alpha) - f(H_0)) \chi_R] = \int_0^\infty \xi_{\alpha,R}(\lambda) f'(\lambda) d\lambda,$$

$$\xi_{\alpha,R}(\lambda) = -F_\alpha(\sqrt{\lambda}R) - F_{1-\alpha}(\sqrt{\lambda}R) + F_0(\sqrt{\lambda}R) + F_1(\sqrt{\lambda}R),$$

$$\begin{aligned} F_\nu(z) = & \frac{z^{2\nu+4}}{8\sqrt{\pi}} {}_2F_3 \left(\nu + 1, \nu + \frac{3}{2}; 2\nu + 2, \nu + 3, \nu + 3; -z^2 \right) \\ & + \frac{z^{2\nu+2}}{4\sqrt{\pi}} {}_2F_3 \left(\nu + \frac{1}{2}, \nu + 1; 2\nu + 1, \nu + 2, \nu + 2; -z^2 \right). \end{aligned}$$

Outline of Proof

Combining Proposition 2 and the asymptotic formula for ${}_2F_3$, we obtain more detailed asymptotics of the function $\xi_{\alpha,R}$ as follows.

Proposition 3

Let $0 < \alpha < 1$. Then we have

$$\begin{aligned}\xi_{\alpha,R}(\lambda) &= \frac{1}{2}\alpha(1-\alpha) - \frac{\sin \alpha\pi}{4\pi} \frac{\cos(2\sqrt{\lambda}R)}{\sqrt{\lambda}R} \\ &\quad + \frac{(2\alpha+1)(2\alpha-3)}{16\pi} \frac{\sin \alpha\pi}{(\sqrt{\lambda}R)^2} \frac{\sin(2\sqrt{\lambda}R)}{(\sqrt{\lambda}R)^2} + O((\sqrt{\lambda}R)^{-3}),\end{aligned}$$

as $\sqrt{\lambda}R \rightarrow \infty$.

The principal term coincides with Tamura's one, but the next term differs because of the difference of the formulation (Tamura uses the smooth cut-off).